## Геология планет

А.Т. Базилевский Институт геохимии и аналитической химии (ГЕОХИ) им. В.И. Вернадского Российской академии наук Москва, Россия



Краткий курс лекций ГЕОХИ РАН Март - Май, 2016

#### The course content

- 1. Solar system, Earth
- 2. The Moon and Mercury
- 3. Venus
- 4. Mars
- 5. Small bodies: Asteroids, comets, meteorites
- 6. Outer planets and their satellites
- 7. Geological processes in the Solar system (summary)

# Solar system

### Our Sun is a star of spectral class G2V

Distance from Earth
Diameter of the Sun
Mass of the Sun

Mean density
Density in the center

Temperature of photosphere at the center

Period of rotation, equator poles

Composition (mass. %)





Sun in X-ray

149 million km = 1 AU 1.4 million km

333,000 mass of Earth

1.4 g/cm<sup>3</sup>
160 g/cm<sup>3</sup>

50lar Spectrum 4300 - 4400 Angstroms

6,000 K

15,000,000

4300

25.6 days

36 dfays

H 73.46 Fe 0.16 M 0.05

He 24.85 Ne 0.12 S 0.04

O 0.77 N 0.09

C 0.29 Si 0.07 other 0.1

- Solar wind
- 11-year cycle of magnetic activity

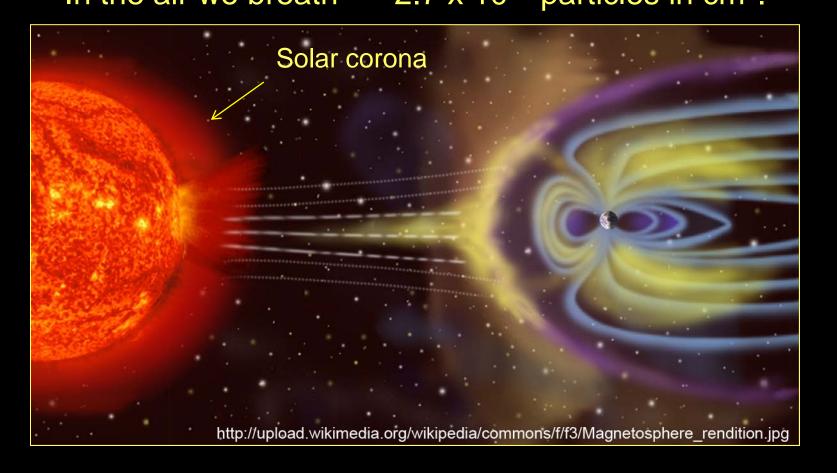
# Interplanetary matter and magnitospheres

- Finely disseminated matter between planets and other bodies of the Solar system системы.
- Includes dust, cosmic rays and hot plasma of solar wind. Temperature of interplanetary matter ~ 100,000 K, while its density in the vicinity of Earth ~ 5 particles / cm³; it decreases with increase of distance from the Sun as R-².
- Plasma caries magnetic field of the Sun.
- Outer edge of the Solar system is a boundary between flux of solar wind and interstar matter – so called heliopause (~100 AU from the Sun).

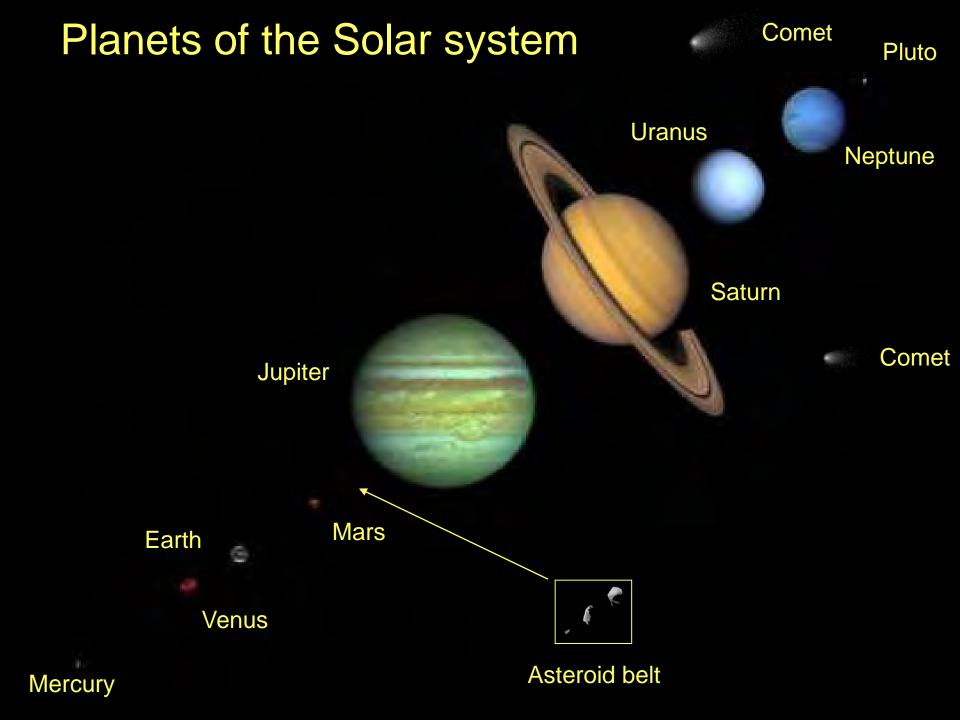
Launched in 1977 Voyager 1 is now in "Interstellar space" (134 AU) and Voyager 2 is currently (111 AU) in the outermost layer of the heliosphere where the solar wind is slowed by the pressure of interstellar gas.

Solar wind - particles, mostly, electrons and protons, leaving solar corona. Very rarified: in the vicinity of Earth – a few particles per cm<sup>3</sup>.

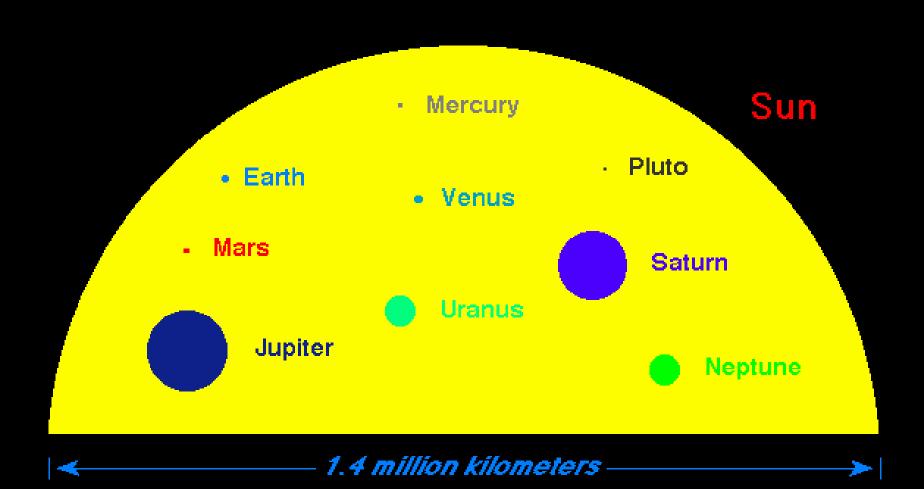
In the air we breath - ~2.7 x 10<sup>19</sup> particles in cm<sup>3</sup>.



Magnetosphere: Part of space, in which magnetic field of planet predominates over magnetic field of solar wind.



## Relative sizes of the Sun and planets



# Planets of the Solar system

Jupiter

Saturn

Uranus

Neptune

8

Pluto

Mars

Mercury

Number

of

known

satelites

Venus

0

0

Earth

	1	2	3	4	5	6	7	8	9
Distance from the Sun, au	0.4	0.7	1.0	1.5	5.2	9.5	19.2	30	39
Radius, km	2,400	6,050	6,378	3,390	71,400	60,300	26,200	25,200	1,200
Mass, M <sub>Earth</sub>	0.06	0.82	1	0.11	317	95	14.5	17	0.003
Density. g/cm <sup>3</sup>	5.4	5.2	5.5	3.9	1.3	0.69	1.3	1.6	2.0
Albedo	0.06	0.76	0.36	0.16	0.43	0.61	0.35	0.35	0.38

16

~60

2

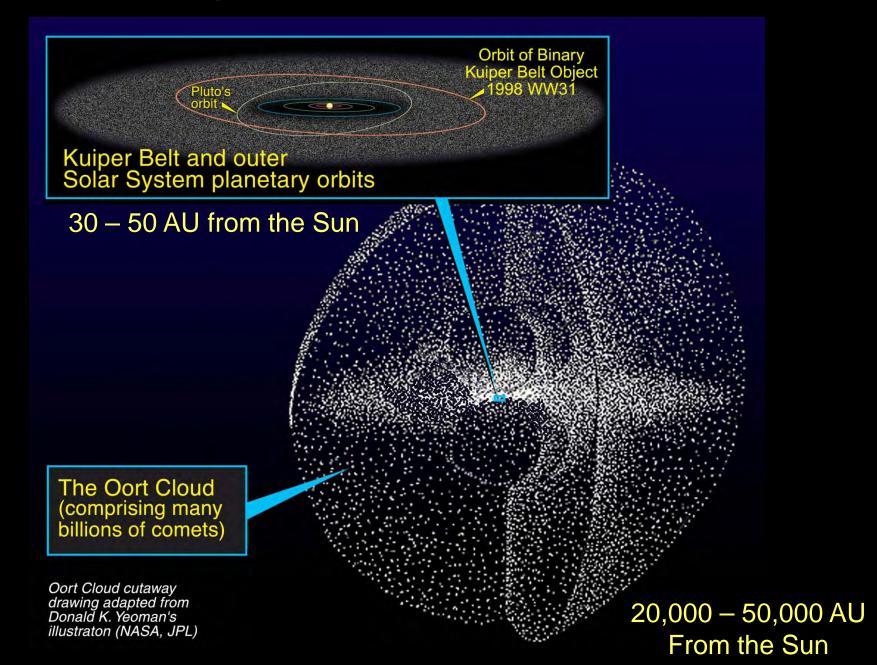
17

~30

15

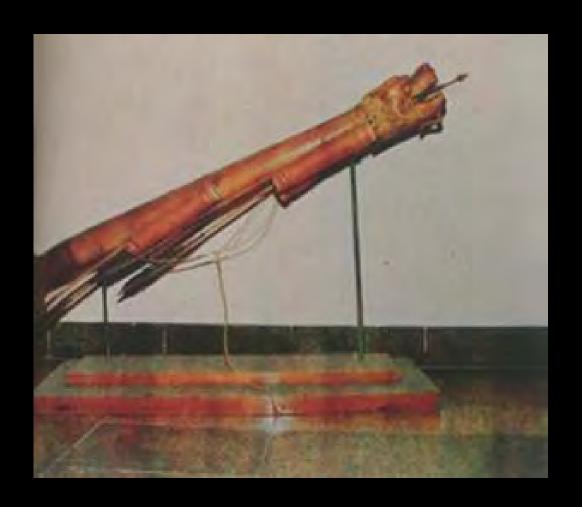
~20

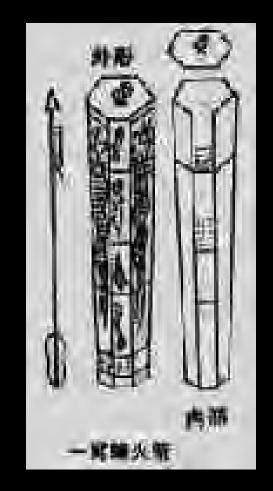
# Kuiper belt and Oort cloud



# Solar system studies became effective with development of rockets

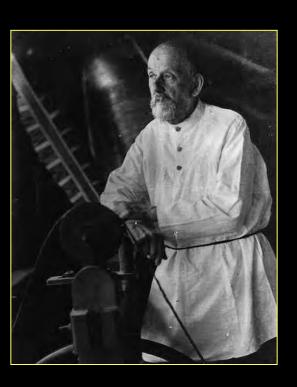
#### Rockets have been invented in ancient China





They were used for military purposes and for fireworks

#### Fathers of space flights





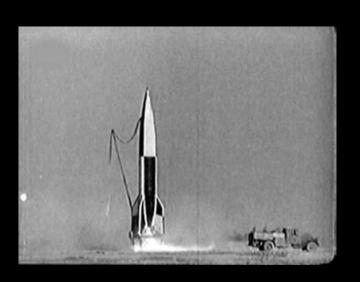


Konstantin Tsiolkovsky
Russian,
Theoretician
of rocketry
1857-1935

Robert
Goddard
American,
Rocket engineer
1882-1945

Hermann
Oberth
German
Rocket engineer
1894-1989

# Space flights to the Moon and further Role of V2 in modern rocket industry



In the Soviet Union End of 40s



In Germany 1944





In USA End of 40s

# The plant, which produced V2, in Peenemunde, Germany









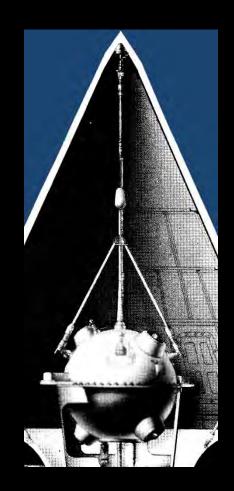
# Beginning of space era: Sputnik and first flights to the Moon



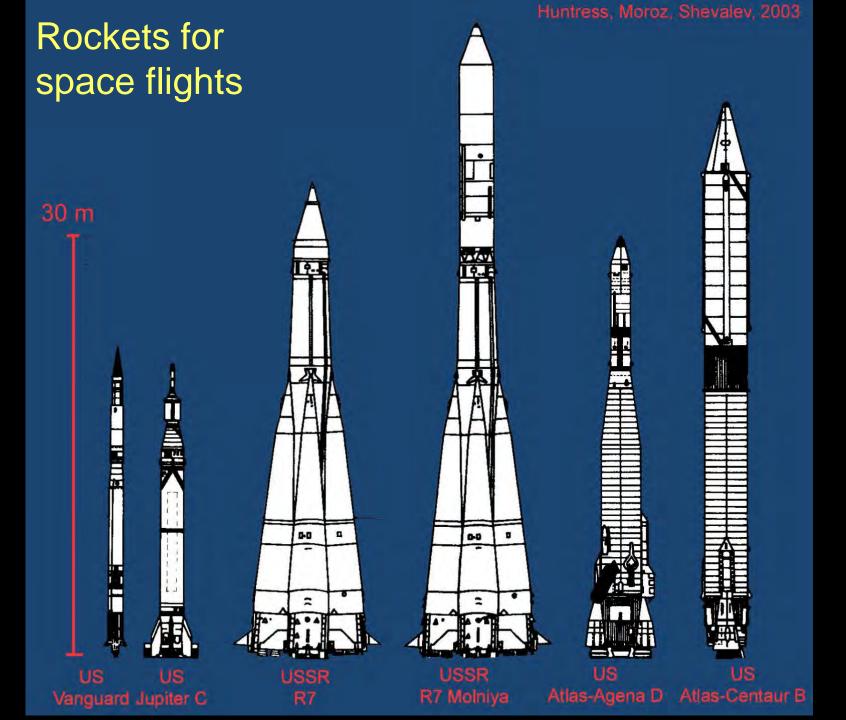
Sputnik October 4, 1957

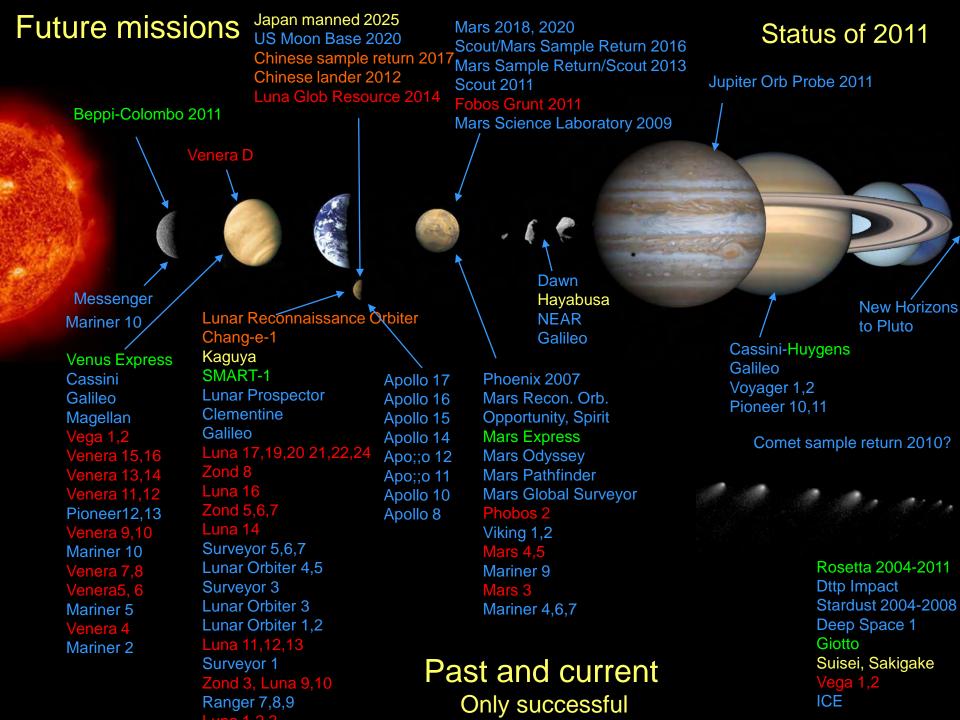


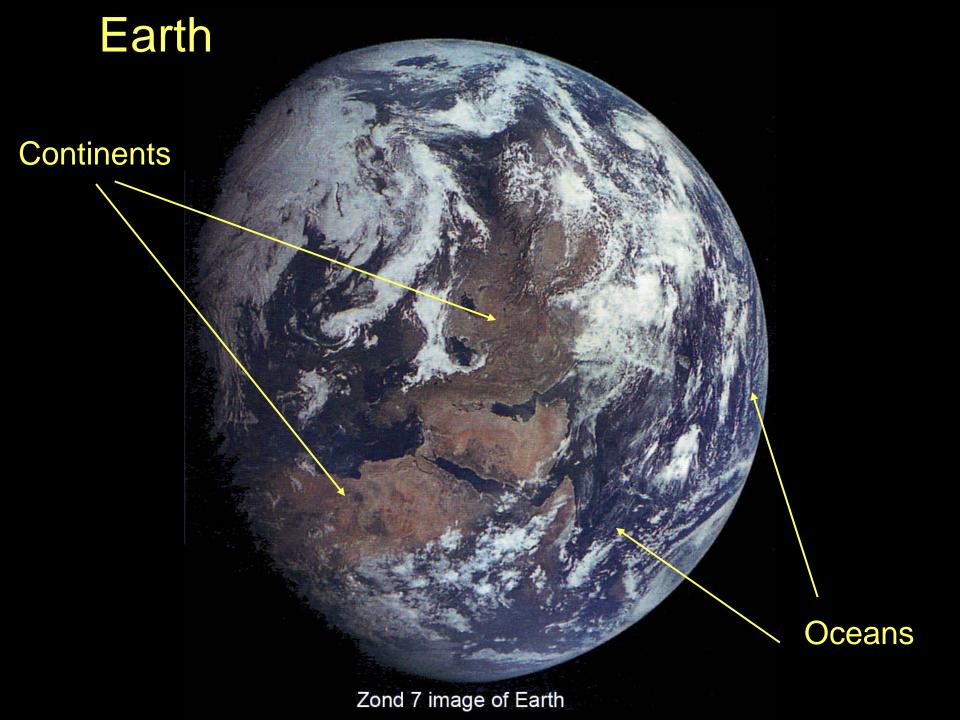
Sergey Korolev Soviet rocket engineer 1906-1966



Luna 1 January 2, 1959







Third planet from the Sun
Distance from the Sun = 149 million km = 1 AU
One satellite – the Moon

Orbital period = 1 year = 365 days Rotation around axis 24 hours

Equatorial diameter 12 756 km Polar diameter 12 712 км

Mass of Earth 5.97 x 10 <sup>24</sup> kg Mean density 5.5 g/cm<sup>3</sup>

Gravity acceleration 978 cm/s<sup>2</sup>

Hydrosphere  $1.35 \times 10^{21} \text{ kg}$ Atmosphere  $5 \times 10^{18} \text{ kg}$ 

> N<sub>2</sub> 78%, O<sub>2</sub> 21%, Ar ~1%, CO<sub>2</sub> 0.04%

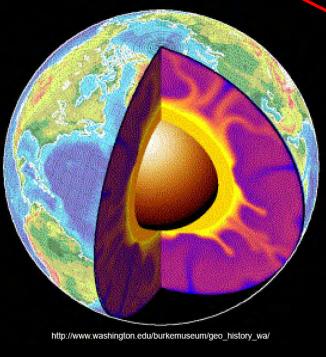


#### Earth interiors

Crust – granites, basalts, sedimentary rocks

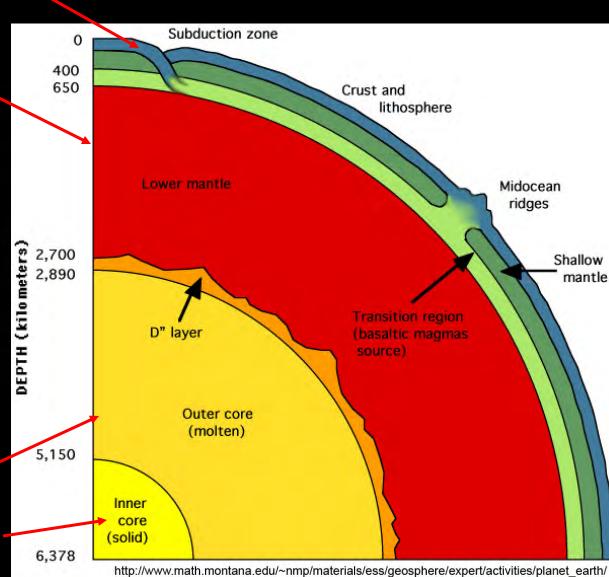
Lithosphere = crust + rigid upper part of the mantle

Mantle- peridotites, dunites

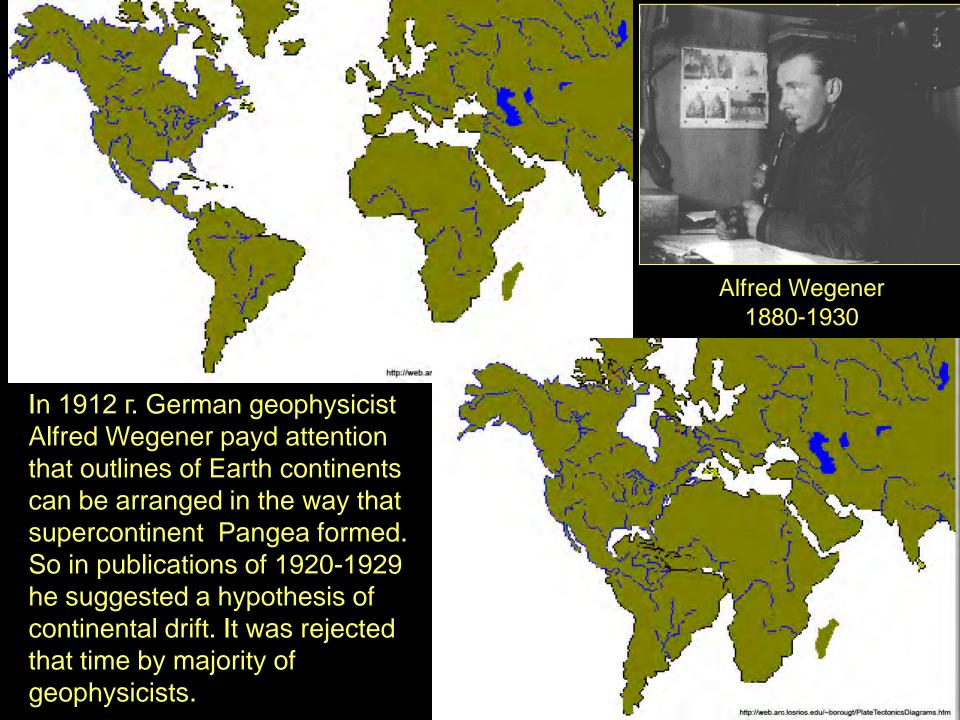


Outer core – melted iron

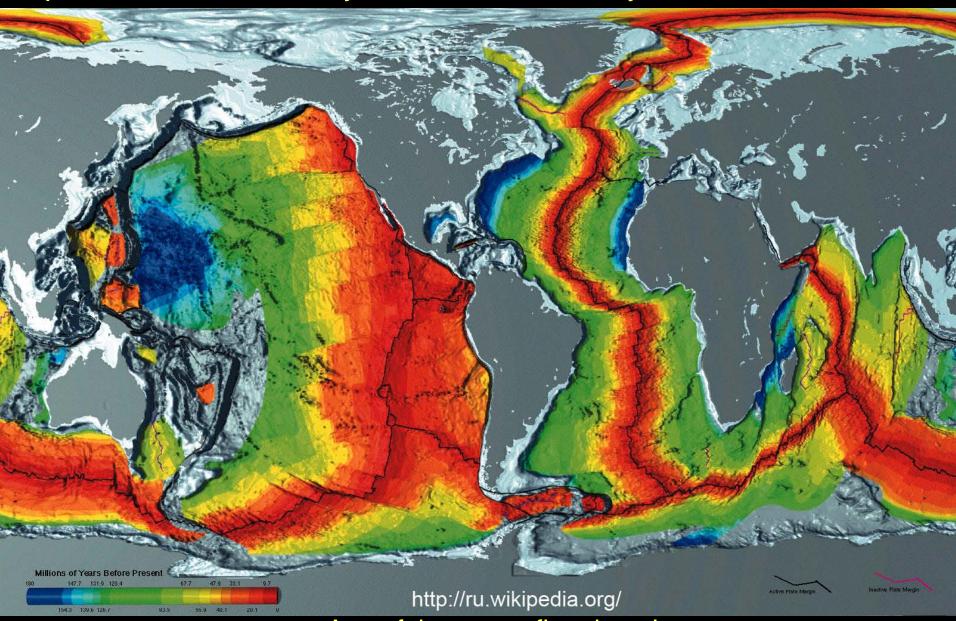
Inner core – solid iron



# Plate tectonics - working heat engine of Earth



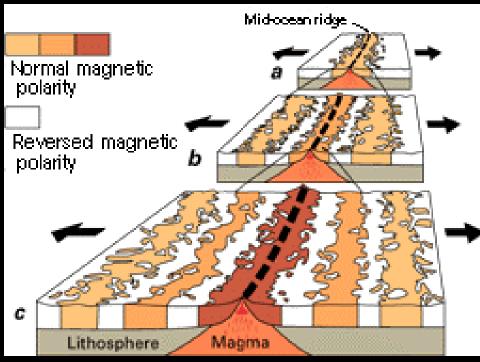
Wegener's hypothesis resurrected in 1960's in the form of plate tectonics, mostly, as a result of study of the ocean floor.



Age of the ocean floor basalts

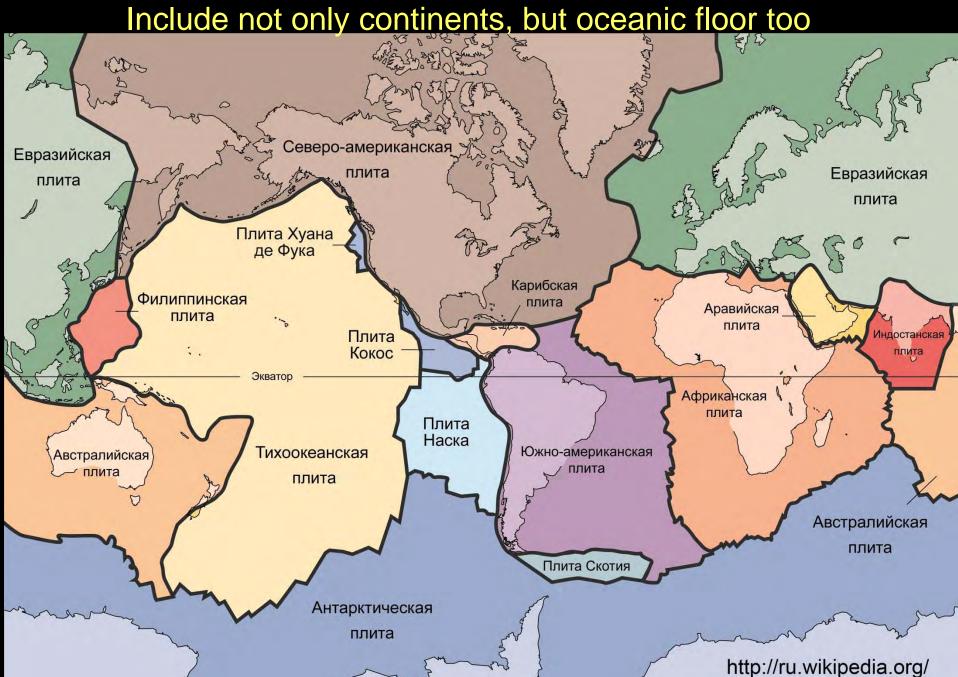
# Queen Charlotte Islands British Columbia Washington Age of oceanic crust (millions of years) Columbia present River Crest of Juan de Fuca Ridge Cape Blanco California Mendocino Crest of Gorda Ridge

# Stripe magnetic anomalies as evidence of oceanic floor spreading



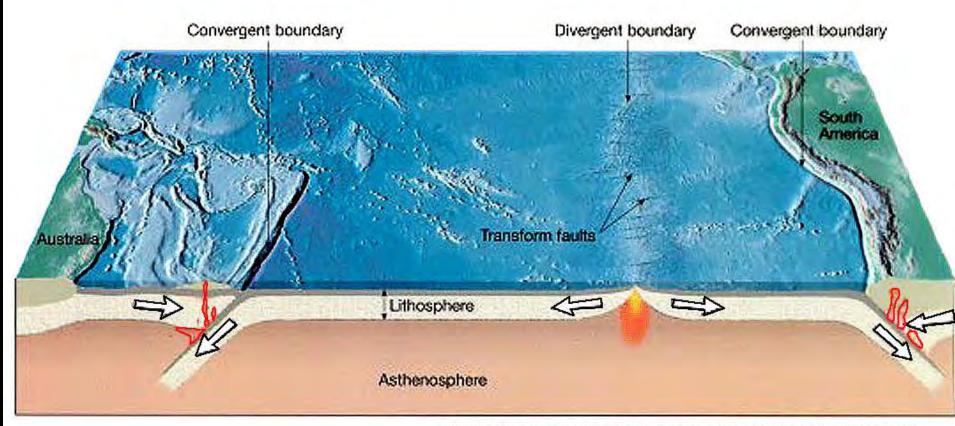
http://web.arc.losrios.edu

Lithospheric pl;ates of Earth.



# Two types of boundaries of lithospheric plates: Divergent and convergent

#### Дивергентные и конвергентные границы плит

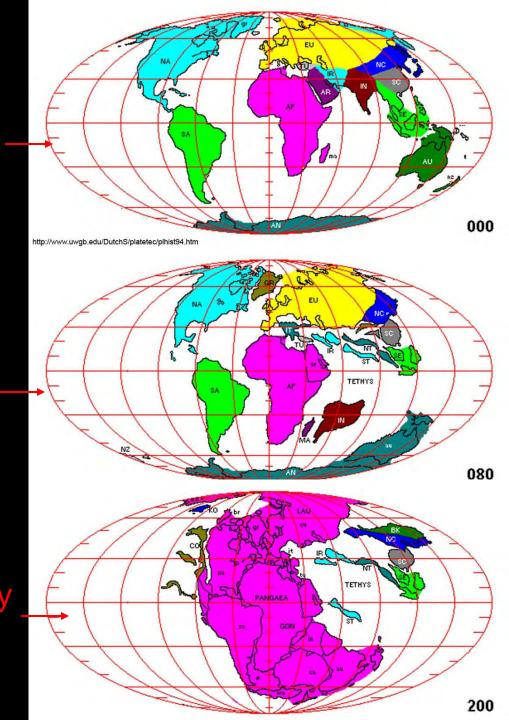


http://web.arc.losrios.edu/~borougt/PlateTectonicsDiagrams.htm

Continents of Earth
now and in the recent
past — Now
— reconstruction
based on geological
and geophysical
data.

In the history of Earth there were 3-4 supercontinents:
2.7, 1.9, 1.2 µ
0.6-0.4 billion years 200 my
ago — ago
supercontinental cycle

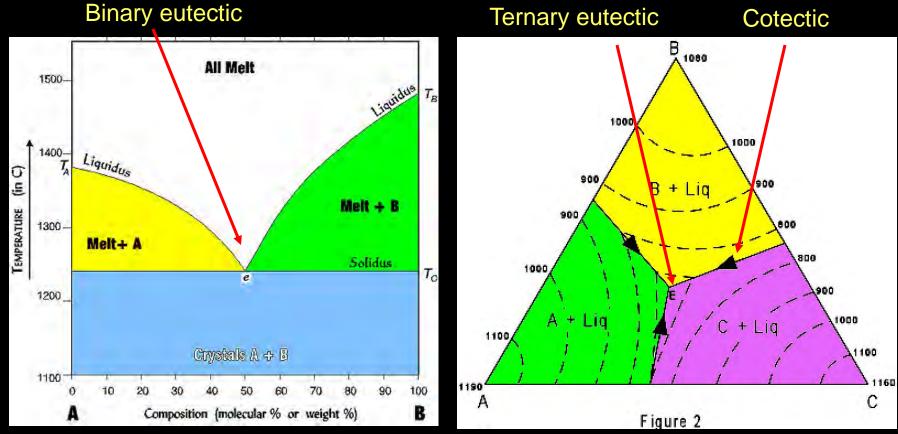
80 my



# Endogenic processes: Volcanism

#### Volcanism

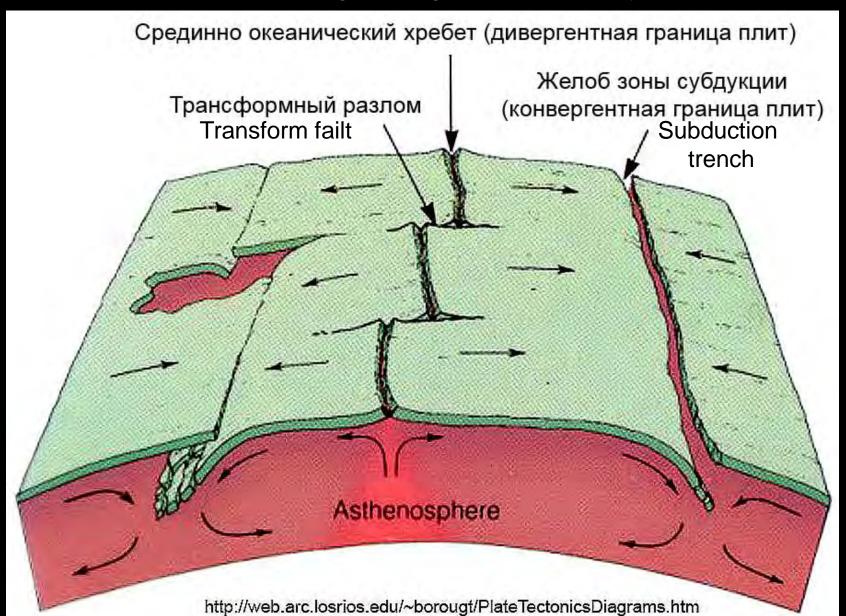
Magmas (melts) appear as the result of partial melting in planet interior, as a rule, due to decrease of pressure in the process of rise of mantle material (convection, hot diapir)



In the process of dry partial melting of the mantle material basaltic magmas formed. In the process of partial melting with participation of H<sub>2</sub>O more differentiated magmas (andesites, granites) formed.

### Volcanism of divergent zones

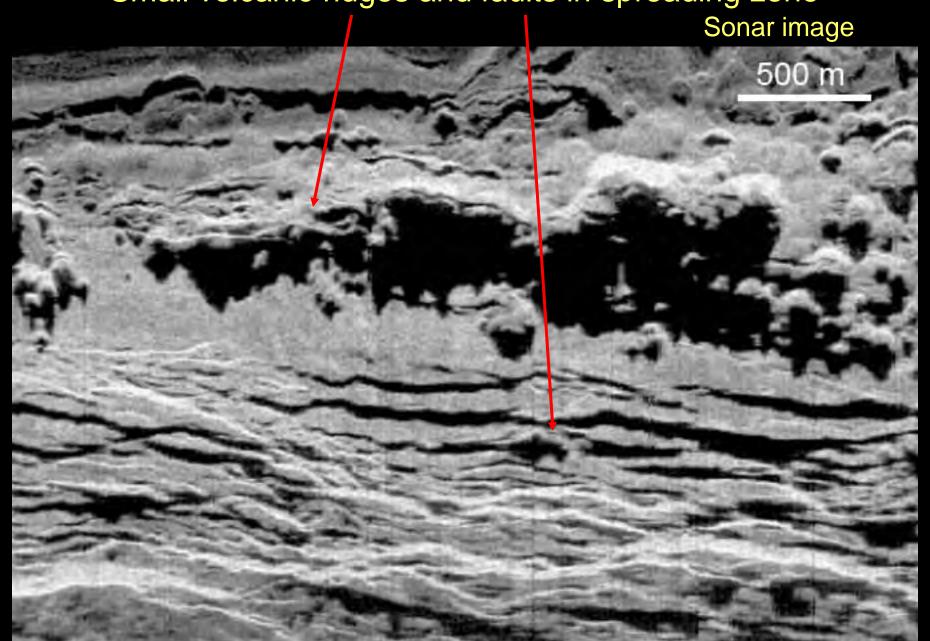
Midoceanic ridge (divergent plate boundary)



#### Eruption of basaltic pillow lavas in spreading zone

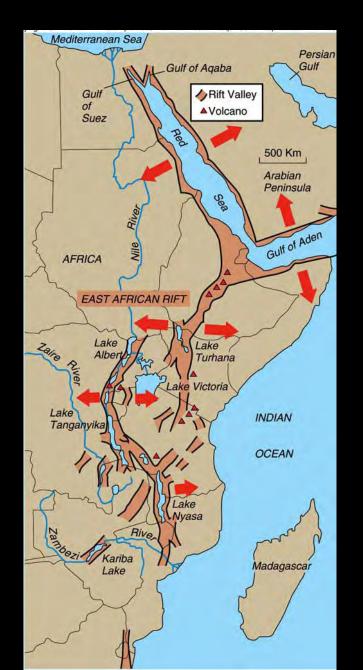


Mid oceanic ridge Small volcanic ridges and faults in spreading zone



#### Rift zones of East Africa

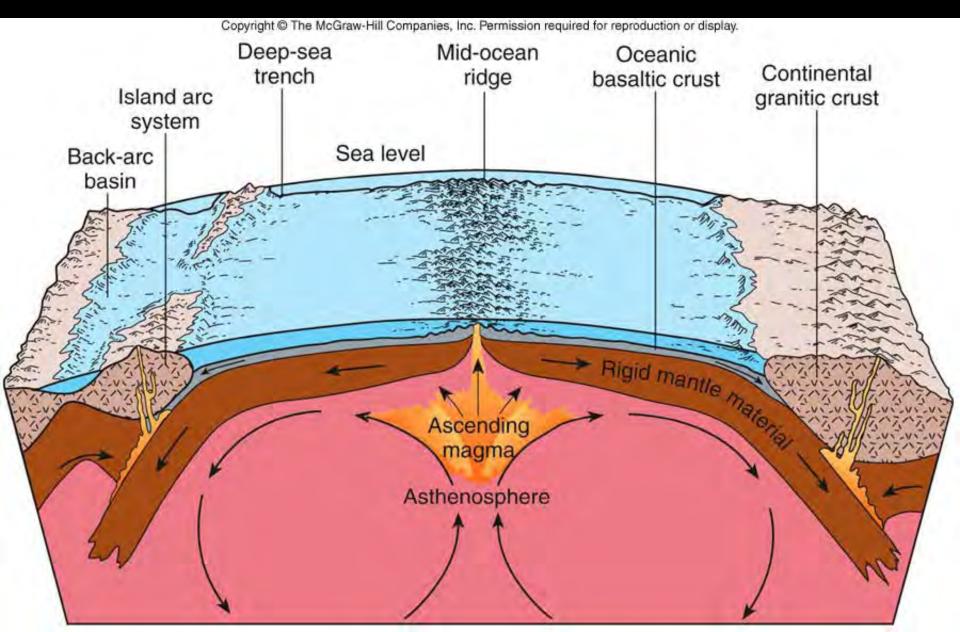




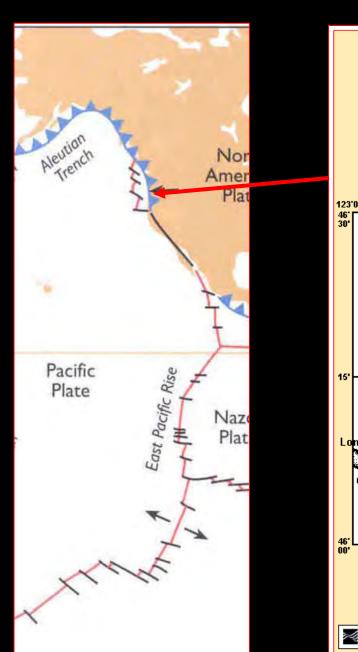
## Nyiragongo volcano in rift zone of East Africa

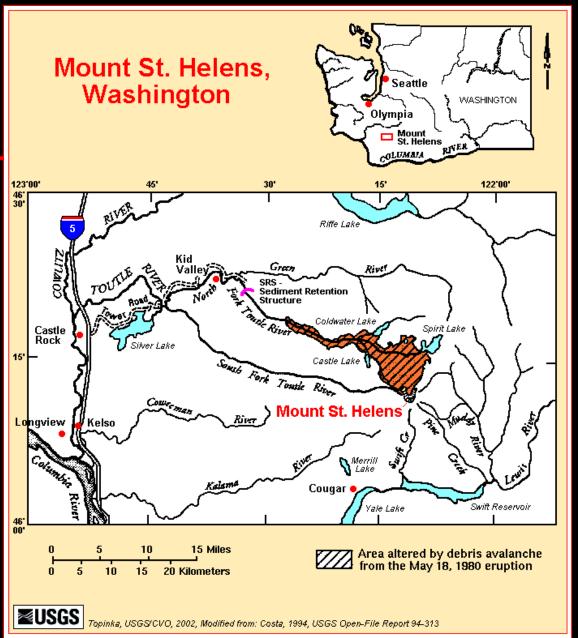


### Divergent and convergent plate boundaries: Convergence case



### St. Helens volcano location

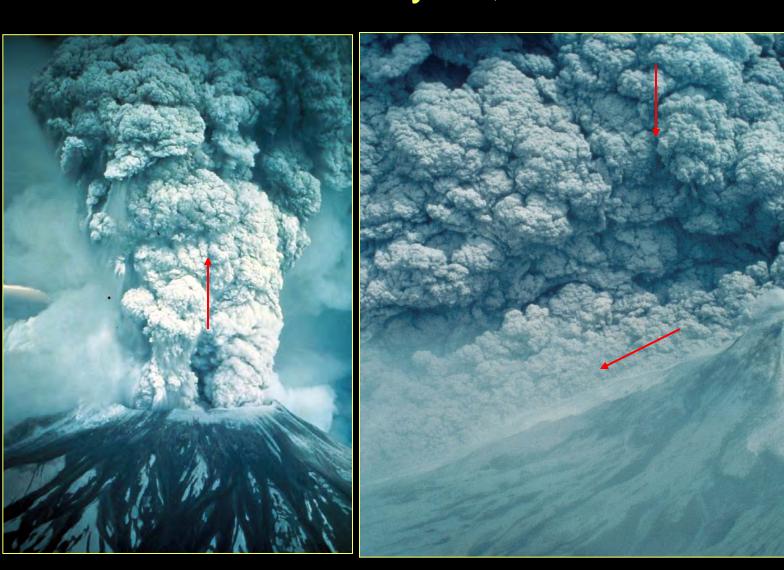




### St. Helens volcano: pre eruption



### St. Helens blast and nuee ardente May 18, 1980



Blast

Nuee ardente

### St. Helens pyroclastic deposit

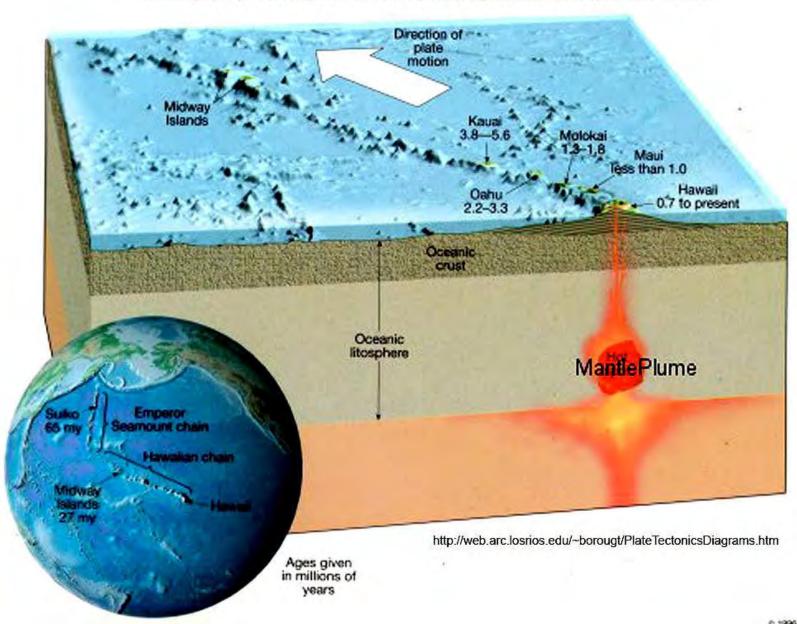


### St. Helens pyroclastic tongue

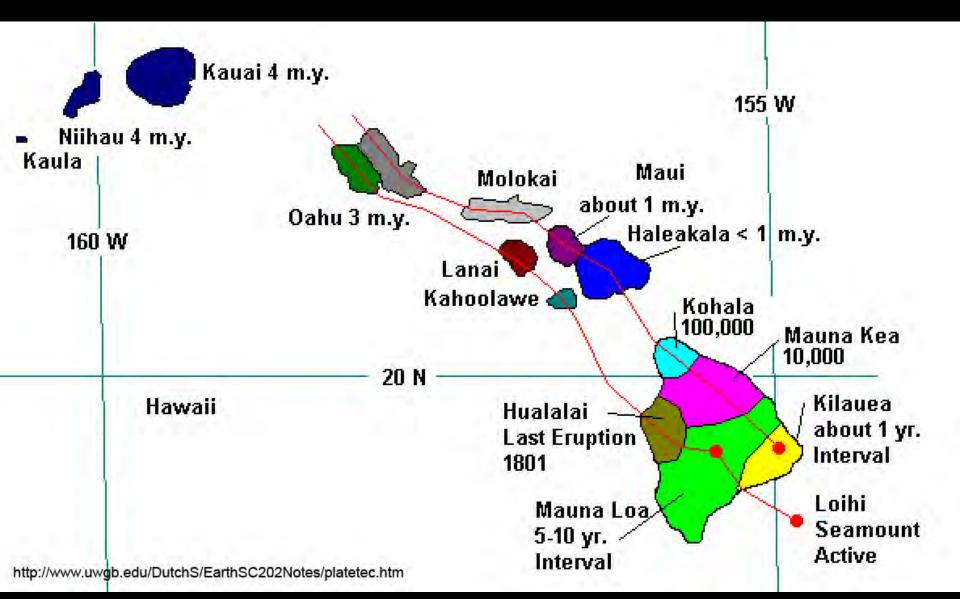


### Hot spot volcanism

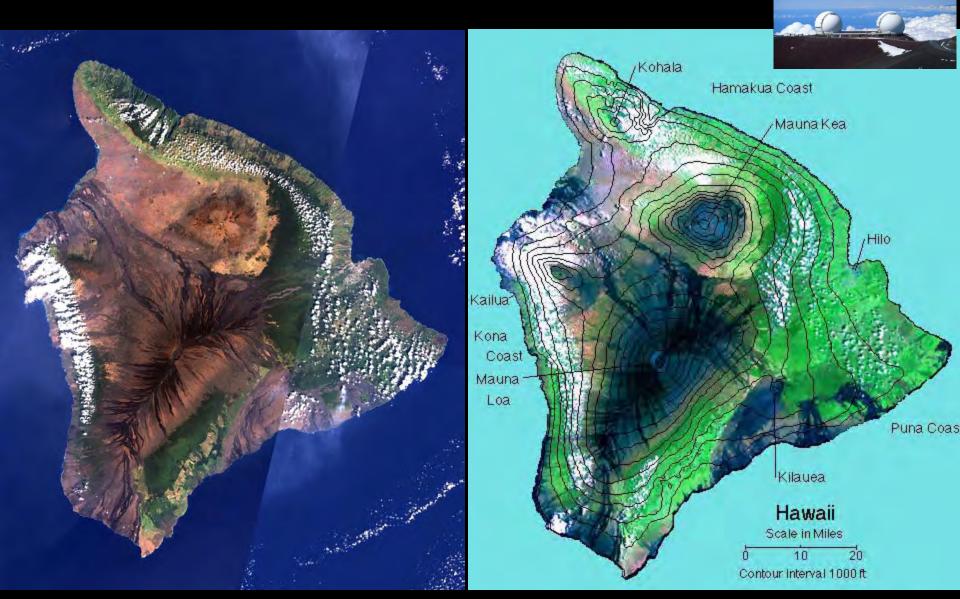
Императорский хребет, Гавайи, и вулканизм горячих точек



### Hot spot volcanism Age of volcanism on the islands increases towards north-east



#### Big island Hawaii: Hot spot volcanism



Landsat image

Mauna Kea

### Hawaii, Summit of Mauna Loa volcano



Photo of USGS

#### Eruption in Kilawea caldera

### Eruption of Mauna Kea volcano



Photos of USGS

# Hawaii, solidified lavas aa and pahoehoe Courtesy of USGS

### Fountaining of lava forms cinder cones, Hawaii



#### Cinder cone and Mauna Kea volcano behind

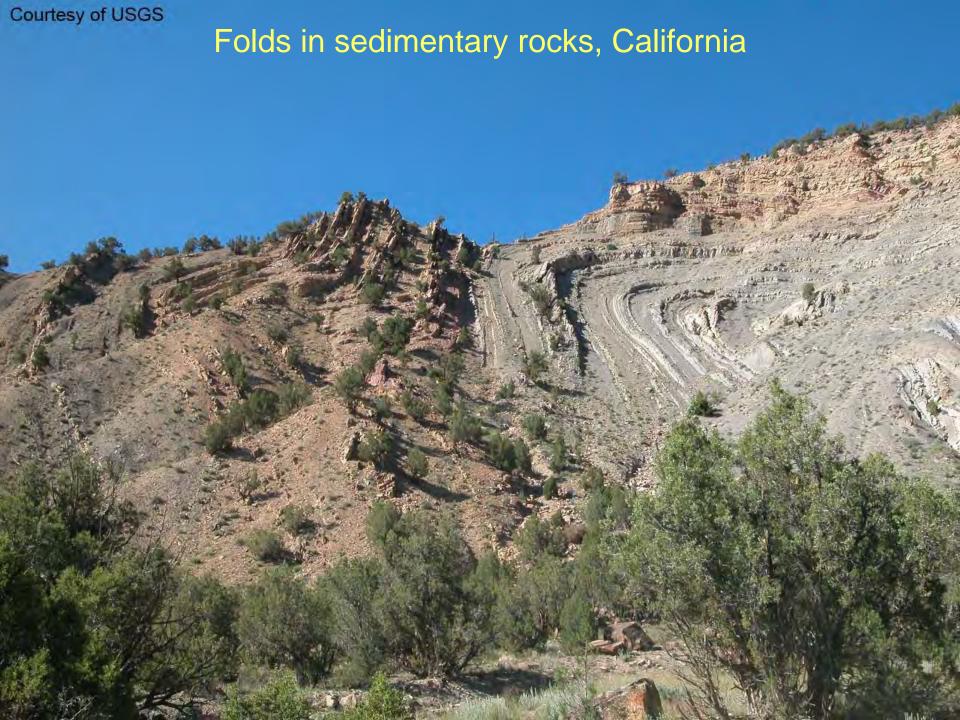


### Endogenic processes: Tectonics –

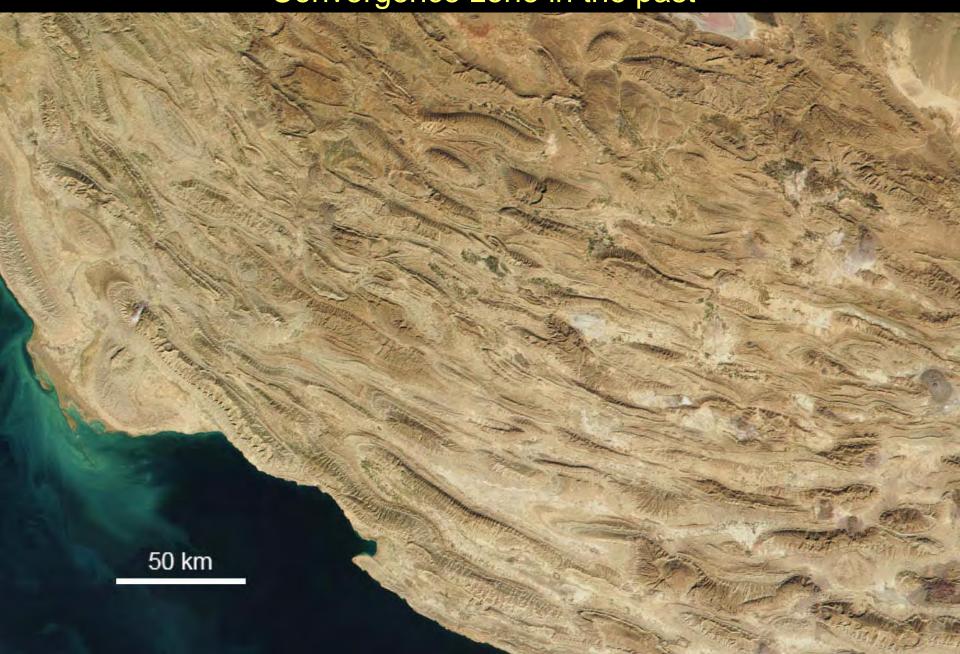
- Deformation of rocks of the Earth crust

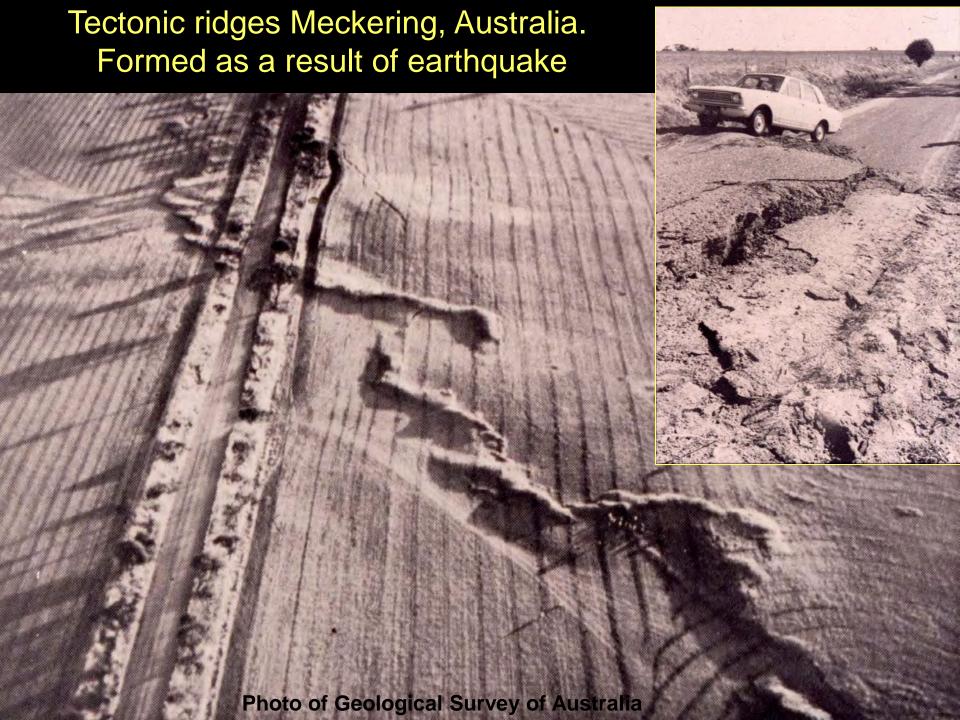
Spreading: Rifts of the Red Sea and Akaba Bay





Folds to the north from Normuz strait: Convergence zone in the past





### Ttectonic breccia Red Point, California



### Exogenic processes: Work of water

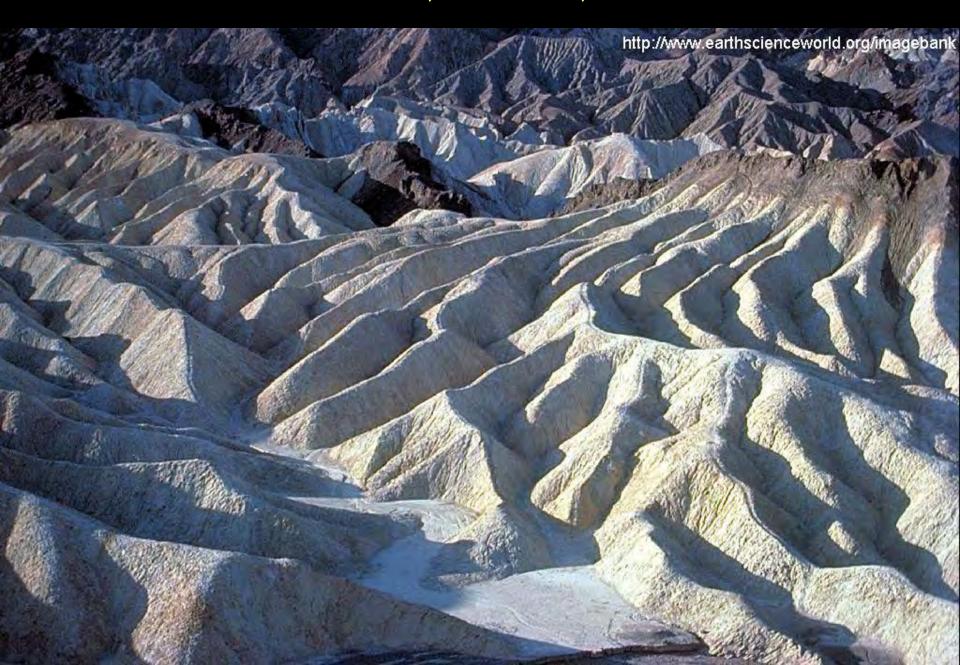
### Shivwits Plateau Grand Canyon of the Colorado River \$ X Courtesy of GeoEye Space Imaging

### Grand Canyon, Colorado, USA





### Badlands, California, USA



### Meanders of Pecatonica river, Wyoming, USA



### Alluvial deposits, Southern Ontario, Canada: Sand, gravel, pebbles



Lacustrine deposits, Southern Ontario, Canada: Silt (very fine sand), clay



### Death Valley, California, USA



### Salt deposits in Death Valley, California, USA



### Rocky shore Point Reies, California, USA



#### Pebble shore of eastern Sussex, England



#### Layers of limestone (CaCO<sub>3</sub>), Arizona, USA



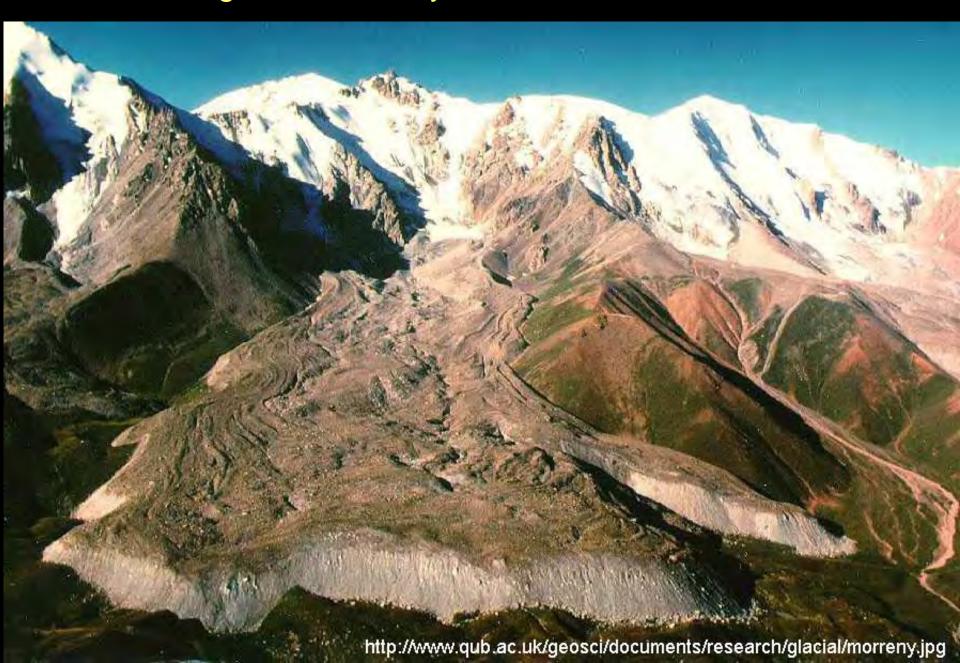
In most cases, limestones are remnants of organisms (shells, etc.)

## Exogenic processes: Work of ice and permafrost phenomena

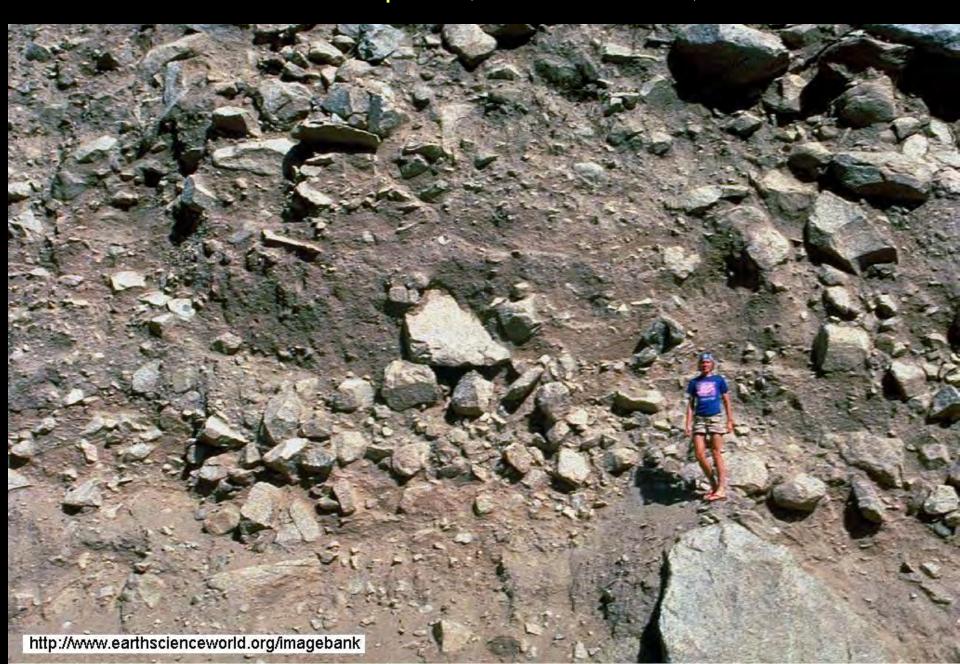
### Atabaska glacier, Alaska, USA



### Rock glacier Morenny, Tien-Shan, Middle Asia



### Moraine deposits, Sierra Nevada, USA



#### Permafrost polygons in North America



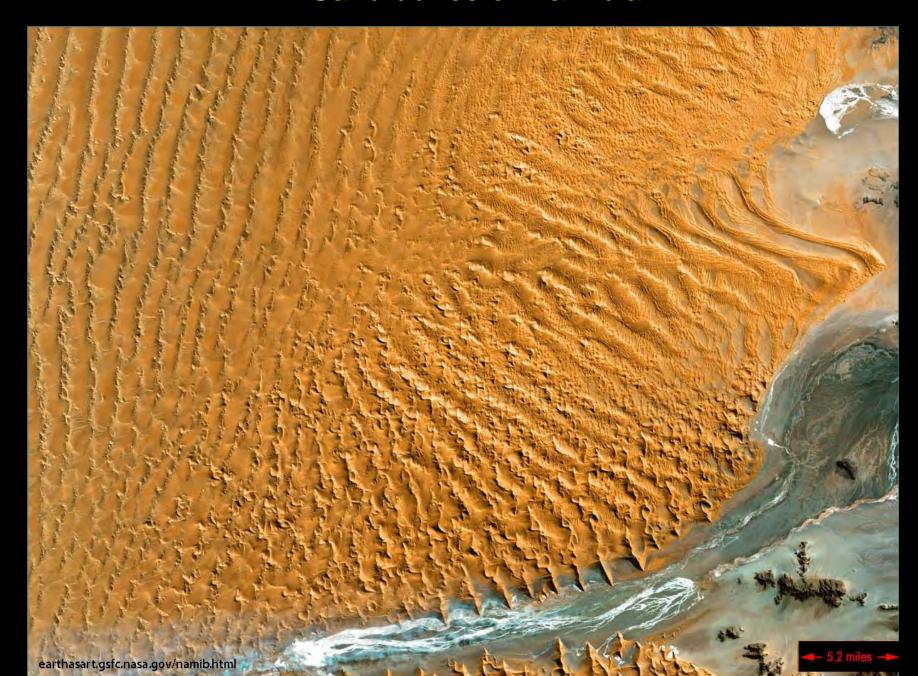
## Exogenic processes: Work of wind

Atlantic coast of Namibia, Africa:

Dust brought by wind into the ocean



#### Sand dunes of Namibia



#### Yardangs - parallel ridges, cut by wind



# Exogenic processes: Work of gravity (Downslope movement)

#### Landslide, Los Angeles, USA



### Down-slided material, Los Angeles, USA Formed in one short in time event. This was provoked by earthquake



### Talus at the foot of a slope, California, USA Accumulated gradually – piece by piece



## Impacts of asteroids, comets and meteorites

#### Meteoritic crater Sterlitamak Russia, 1990



Impact crater Meteor,
Arizona, USA D = 1.2 kmFormed ~50,000 years ago





Craters Clear Water West (36 km) & East (26 km)

Quebek, Canada



## Life and its influence on the planet Earth

### Underwater life Coral reef. Part of limestones – are lithified corals



#### Life on the Earth surface



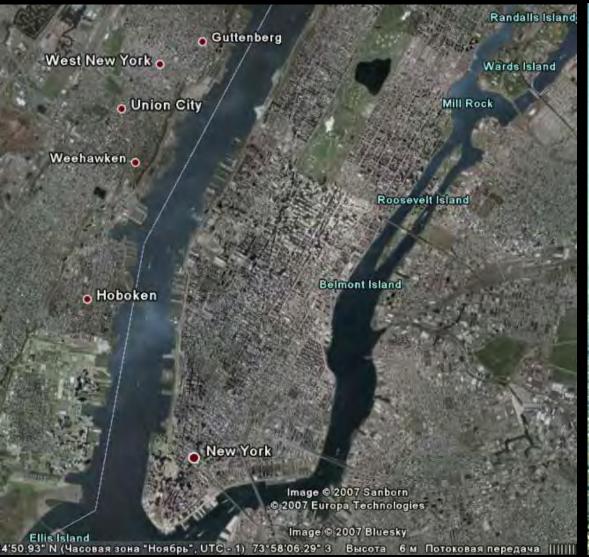
#### People



#### Anthropogenic landscape, Madagascar

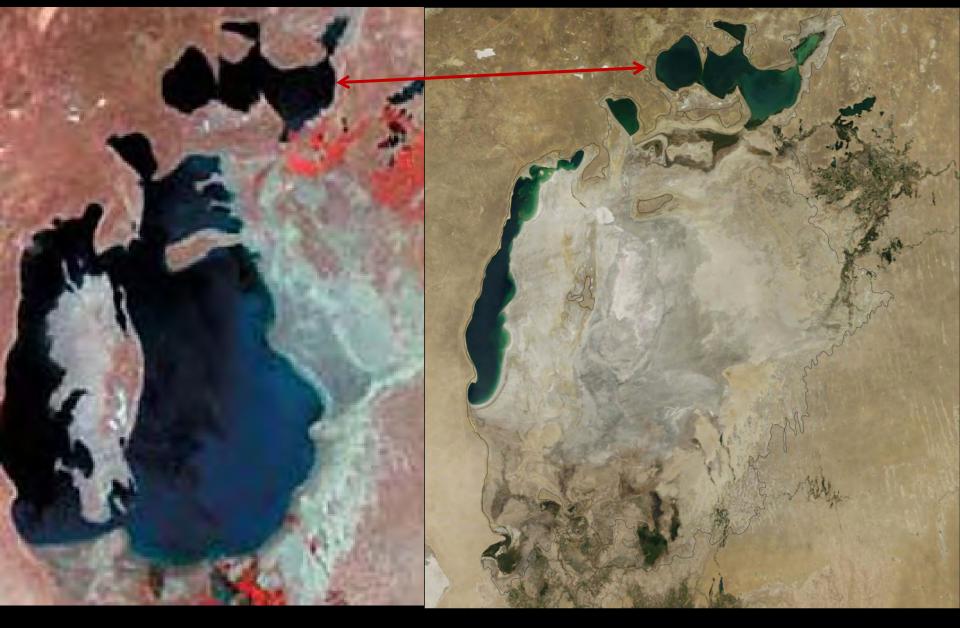


#### Anthropogenic landscape, New York city, USA





Drying of the Aral Sea, Middle Asia – Kazakhstan // Uzbekistan



1988 2015

#### Earth, view from the near-lunar orbit

